



FUV Detector System Performance Verification

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FUV PERFORMANCE REQUIREMENTS



- COS FUV Detector performance requirements given in document COS-08-0003 (Statement of Requirements)
 - Quantum Efficiency both in and out of band
 - Spatial Resolution
 - Spatial Linearity
 - Dark Count Rate
 - Maximum Global and Local Count Rate
 - Deadtime Performance
 - Detector Lifetime
- Performance Verification Test Plan in COS-UCB-006
- Verification done with standard and specialized test procedures throughout detector system build.



FUV Performance Requirements (cont)



Item	Performance Requirement	SOR
Image Quality		
Spatial resolution	$<(25 \mu\text{m} \times 50 \mu\text{m})$ FWHM over 80% of active area	3.7
Integral non-linearity ($\geq 1\text{mm}$ frequency)	$<\pm 100 \mu\text{m}$ over 80% of active area	3.8.1
Differential non-linearity ($<1\text{mm}$ frequency)	Characterize to 3.2% RMS (Flat field at 166 cts /pixel)	3.8.2
Thermal spatial stability	$<\pm 3.3$ pixels/ $^{\circ}\text{C}$ (DVA), ± 1.7 pixels/ $^{\circ}\text{C}$ (DEB)	3.15
Flat field stability	$\leq 1\%$ RMS per resel over ≥ 80 mm by 0.5 mm	3.16
Detection Efficiency		
Quantum efficiency (QDE)	$\geq 25\%$ @ 1335Å, 19% @ 1463Å, 17% @ 1560Å	3.5
Visible light rejection	QDE $< 10^{-6}$ @ 4000Å to 6500Å	3.5.2
Deadtime efficiency loss	10% QDE loss at 10 kcps per segment 60% QDE loss at 100 kcps per segment	3.13
Non Active Area Scatter	No (Ly α) scattered photons from MCP gap	3.6.1
Event Rates		
Dark count rate (at I&T)	< 0.5 cps/cm ²	3.9
Maximum global count rate	$\geq 40,000$ cps/segment	3.11
Maximum local count rate	5 cps/pore over $10^3 \mu\text{m}^2$ (81cps/ $10^3 \mu\text{m}^2$)	3.12
MCP Fluence Life	$> 10^9$ events per mm ²	3.14



FUV PERFORMANCE VERIFICATION



COS Detector Performance Verification Test Matrix

Vallerga		10/30/00		Test Procedure								Facility					
Requirement	Description	Qualified by Analysis	Test	After photocathode	Gain/PHD vs. Voltage	Background rate image	Flat field	Resolution/Linearity Mask	QDE vs I and q	MCP Lifetime test	Deadtime Calibration	Resolution/UV Microscope	Test Chamber	QE Calib. Tank	DVA	Date passed	Report #
SOR 3.2	Pixel Format	X															
SOR 3.3	Digitized pixel Scale		X		X	X	X						X				
SOR 3.4	Wavelength Range	X															
SOR 3.5	Quantum Efficiency		X	X	X	X		X						X			
SOR 3.5.1	Grid Shadowing		X	X			X	X						X			
SOR 3.5.2	Visible Rejection		X	X		X	X	X						X			
SOR 3.6.1	Non-Active Area		X			X	X		X					X			
SOR 3.7	Spatial Resolution		X		X	X		X					X				
SOR 3.8	Spatial Linearity		X		X	X		X					X				
SOR 3.9	Dark Count Rate		X	X	X	X									X		
SOR 3.10.1	Dead spots		X	X			X							X			
SOR 3.10.2	Hot spots		X	X		X	X								X		
SOR 3.11	Max. global rate		X		X		X						X				
SOR 3.12	Max. Local ct. rate		X		X			X					X				
SOR 3.13	Deadtime Performance		X								X		X				
SOR 3.14	MCP Lifetime		X	(Proxy MCPs)						X			X				



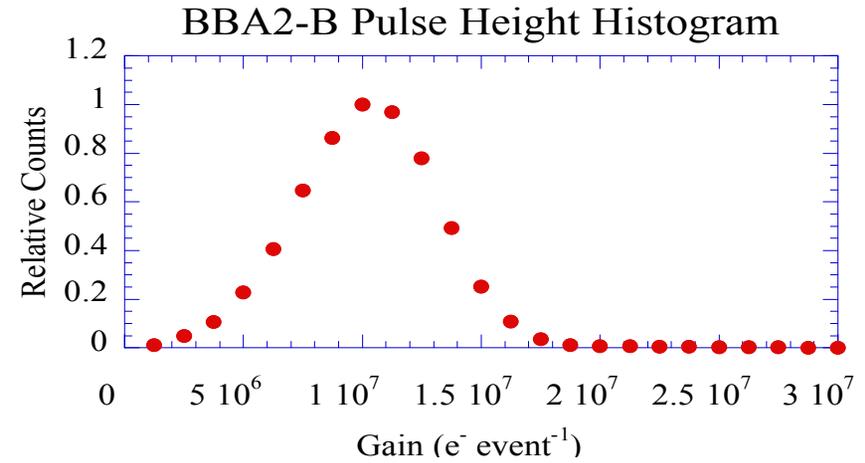
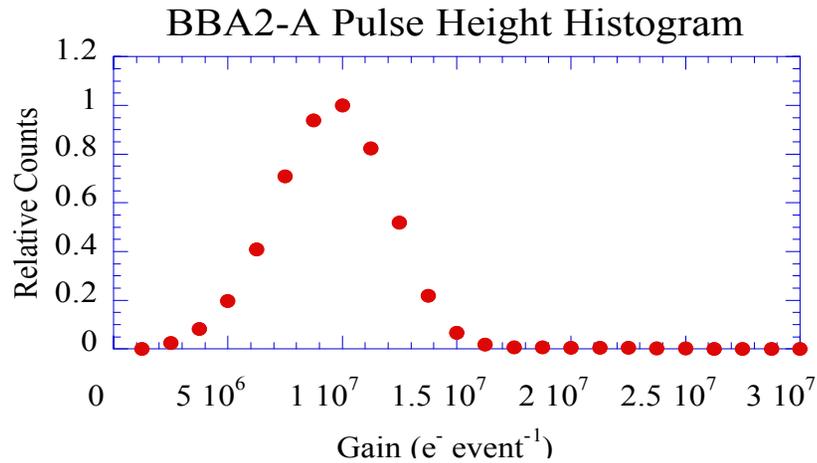
COS Detector Test Flow Plan



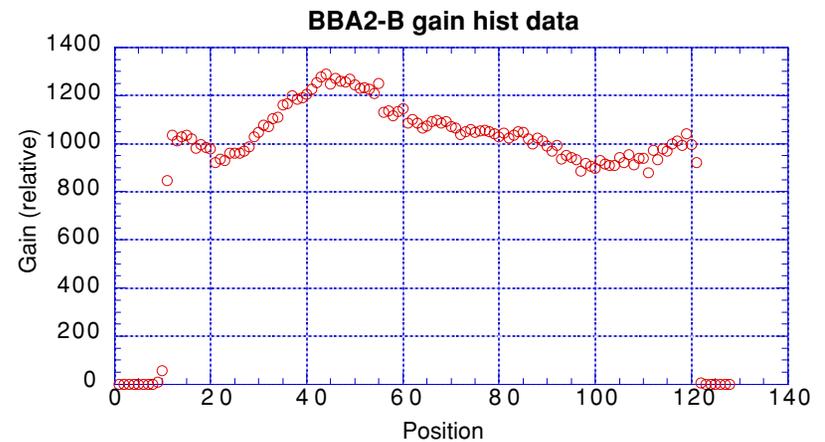
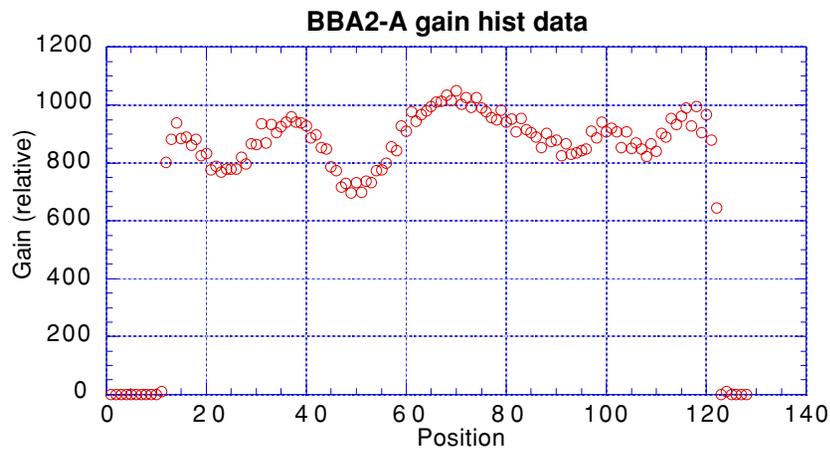
- Select MCP stacks
 - Gain uniformity and PHDs *
 - Background and stability *
 - Flat Field response *
 - QE test (bare) *
 - Assemble Flight Detector
 - Gain uniformity and background *
 - Pinhole Mask tests
 - Set rear field, trim and walk *
 - Resolution test *
 - Spatial linearity test *
 - Local rate test
 - Flat Field tests
 - Global rate test *
 - Deadtime test *
 - Dead spot test *
 - Stability test
 - Deposit Photocathode
 - QE Calibration
 - QE (absolute, angle, position)
 - Visible Rejection test
 - Grid shadow test
 - Non-Active area scattering test
 - Final configuration and install in DVA
 - Environmental tests (Part 1)
 - UV scrub
 - Dark rate test/Hot spots *
 - Final Flight Config. Flat Field
 - Duration TBD
 - Environmental tests (Part 2)
- * Denotes tests which have already demonstrated performance using combinations of Flight and ETU components



Gain Uniformity



PHDs are all $\sim 40\text{-}60\%$ FWHM.

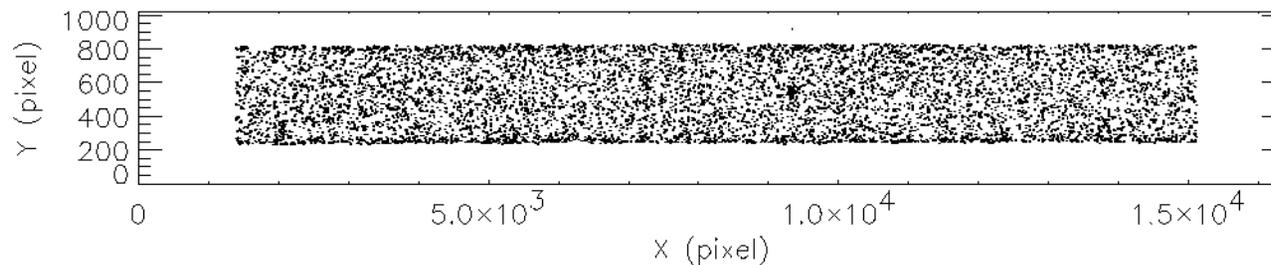




Dark Count Rate / Localized Anomalies



- Requirement of $<0.5 \text{ cts cm}^{-2} \text{ s}^{-1}$ on the ground (not in-orbit)
- Measured throughout testing, both before and after photocathode, without photon or particle illumination
- Localized Anomalies (hot spots or dead spots) shall not compromise more than 2% of spectrum
- COS P.I. can accept or reject plates during selection process
- Most of the COS MCP stacks have $1\text{-}4 \text{ cts s}^{-1} \text{ seg}^{-1}$ ($0.1\text{-}0.5 \text{ cts cm}^{-2} \text{ s}^{-1}$)





Spatial Resolution



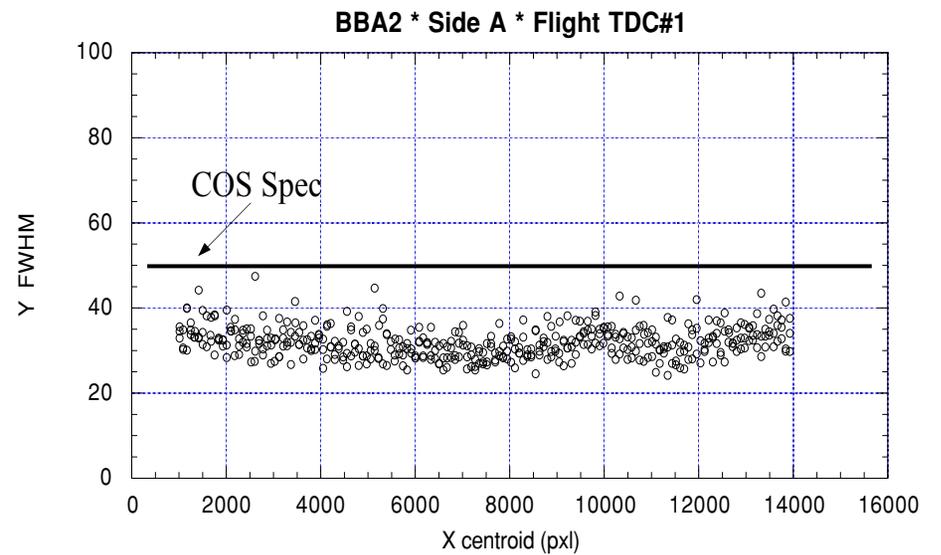
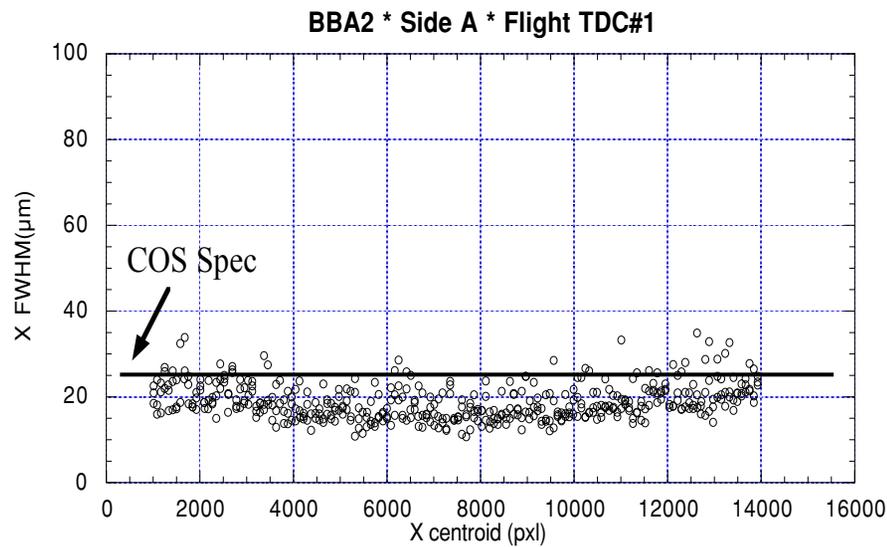
- Pinhole mask (0.5 mm x 0.5 mm grid of 10 μ m pinholes) placed in contact with top MCP and used to sample spatial resolution over whole detector
- Photon list exposure taken at full resolution to get \sim 1000 events per pinhole for \sim 1600 pinholes
- Each spot fit with gaussian and FWHM determined. Systematic scatter in FWHM results is due to illumination of 1, 2 or 3 MCP pores with 10 μ m pinholes. Slit mask used in Linearity Tests (see below) can reduce scatter with loss of resolution
- Trimmed flight electronics required for this test
- A second test uses a slit mask (25 μ m x 500 μ m slits on 200 μ m centers) to measure the X resolution. Also useful for linearity testing.



Spatial Resolution (cont'd)



- Test Mask Sub-Image:

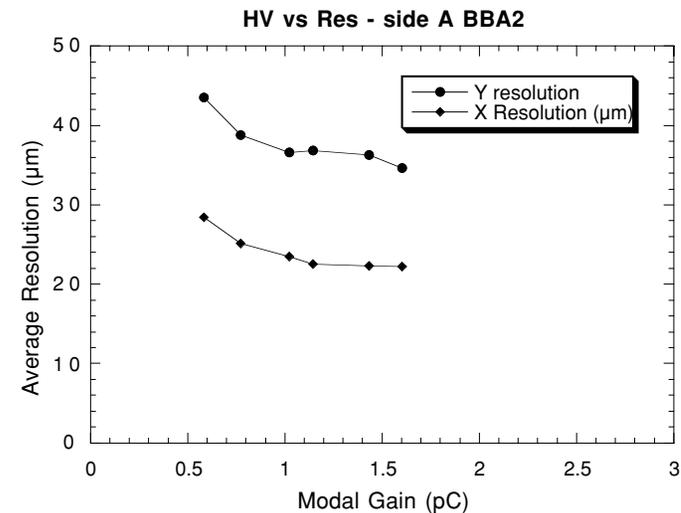
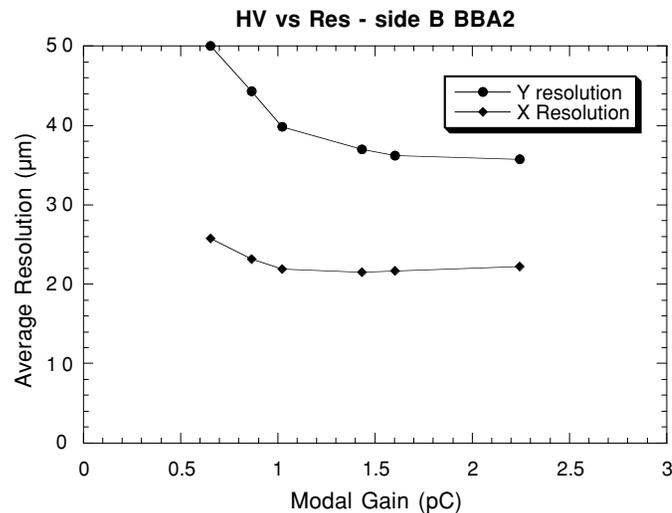
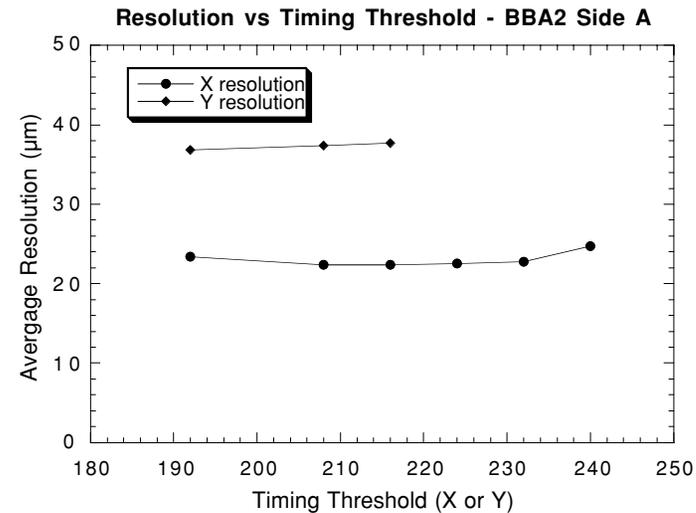
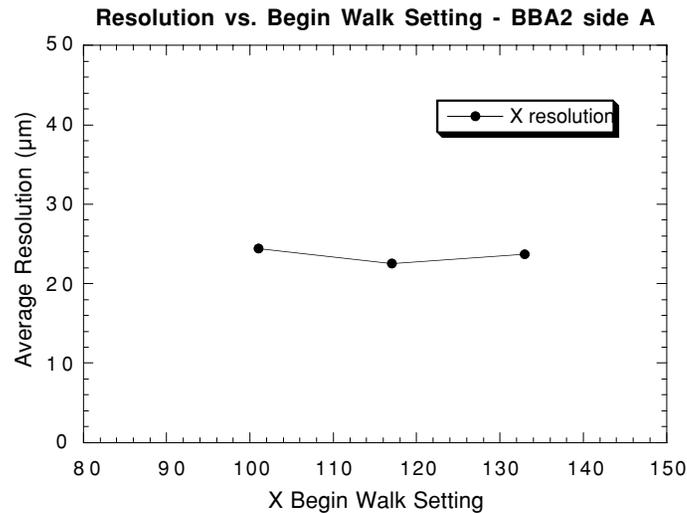




BBA#2 + flight anodes, resolution stability



- Resolution very stable to electronics settings, and over wide gain range





Spatial Linearity

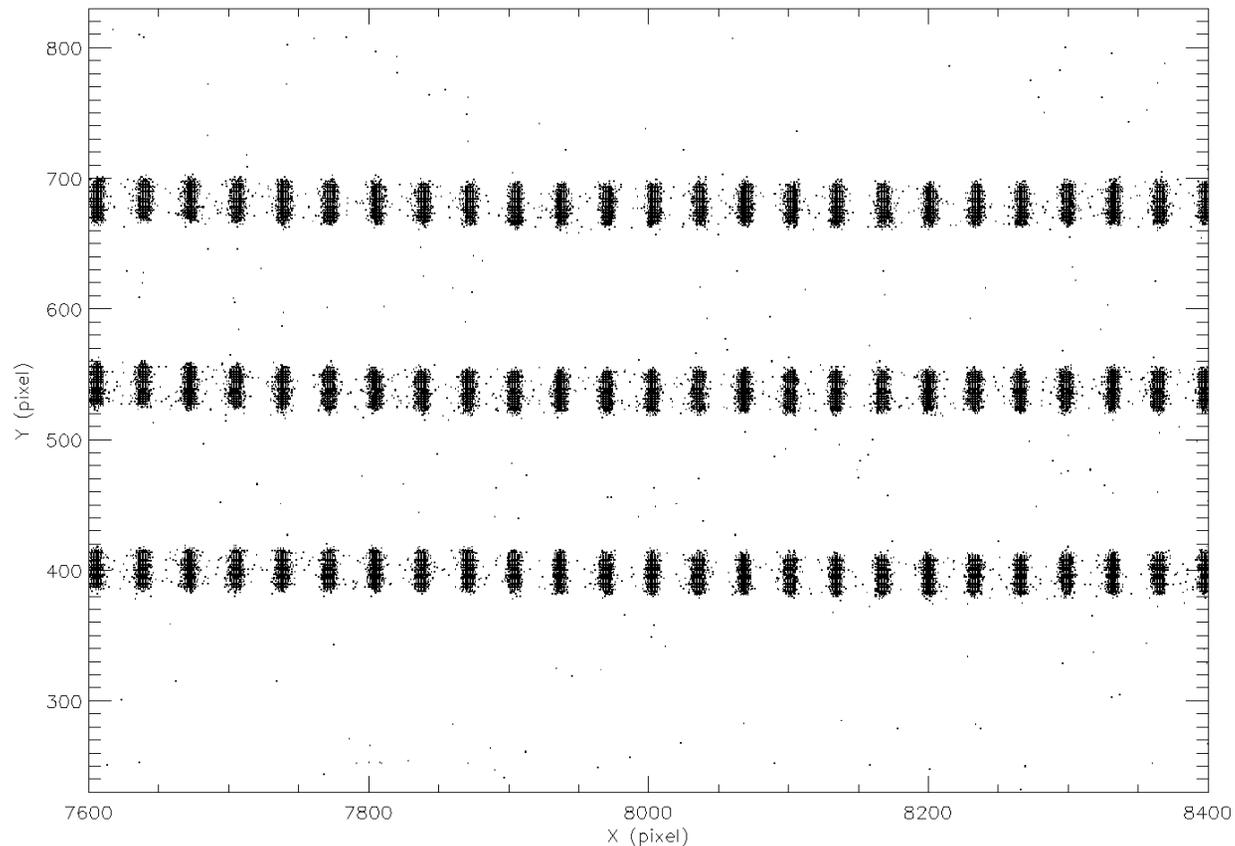


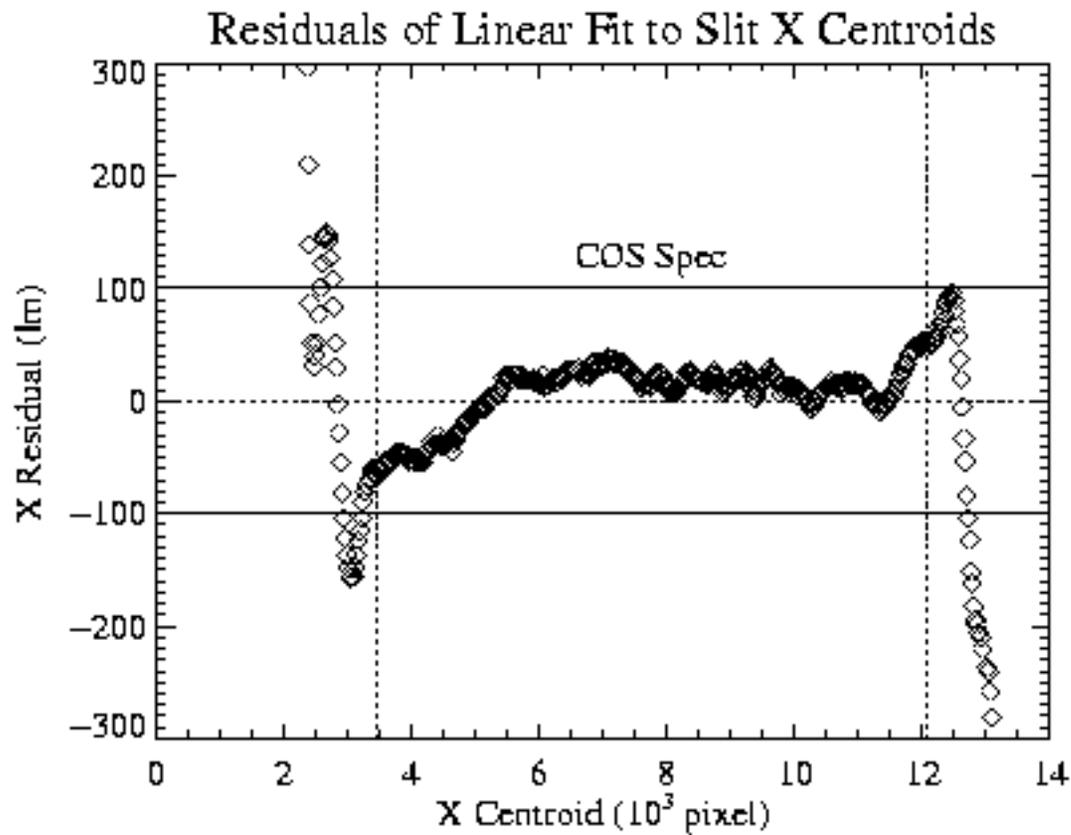
- XDL anodes have “analog” pixels; $X(\text{cm}) \Rightarrow X(\text{pixel})$ to first order is a linear mapping dependent on MCPs, anodes and electronics
- Non-linearity can result in incorrect wavelength and flux density
- Separate into two types requiring different measuring techniques
 - “**Integral**” or slowly varying in spatial dimension ($> 1 \text{ mm}$)
 - “**Differential**” at the highest spatial frequencies
- Requirement
 - Integral non-linearity (true location-measured) shall not exceed $100\mu\text{m}$ and be measured every 1 mm over 80% of the active area.
 - Differential must be characterized by a uniform flat field with 10000 cts/resel to achieve a S/N ratio of $> 100:1$ per resel ($35\mu\text{m} \times 200\mu\text{m}$).
 - This corresponds to a flat field of ~ 1 billion cts or 8 hours at 40,000 cps



Spatial Linearity (cont'd)

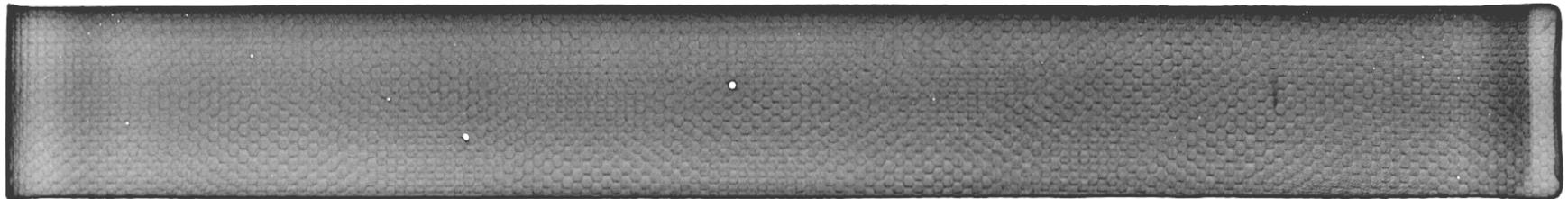
- Use “slit mask”: 3 rows of $25\mu\text{m} \times 500\mu\text{m}$ slits spaced on $200\mu\text{m}$ centers
- Oversamples > 1 mm non-linearities by a factor of 5





$\pm 100 \mu\text{m}$ achieved
over $> 80\%$ of the
detector

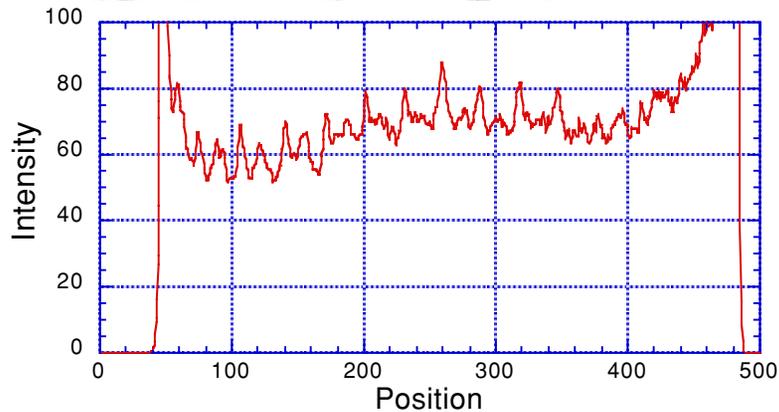
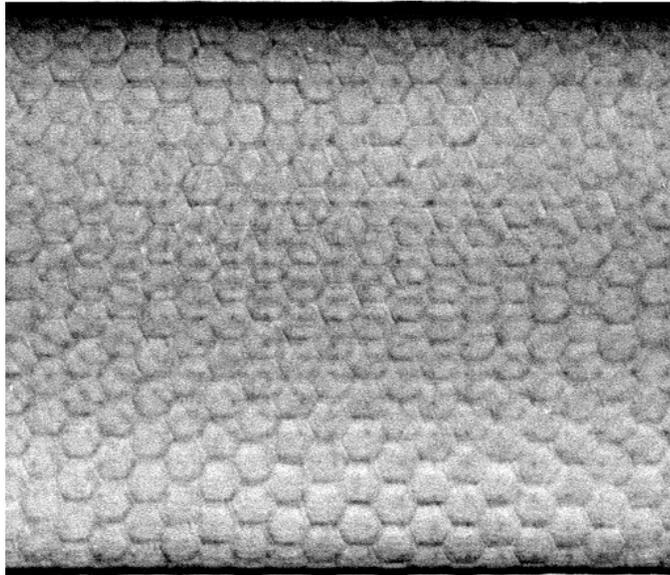
- Small scale distortions cannot be sampled by pinhole grids
 - Multi-fiber boundaries
 - Moiré
- Distortions cause fixed pattern noise since pixels have different effective areas
- Deep, uniform flat fields can characterize pixel size



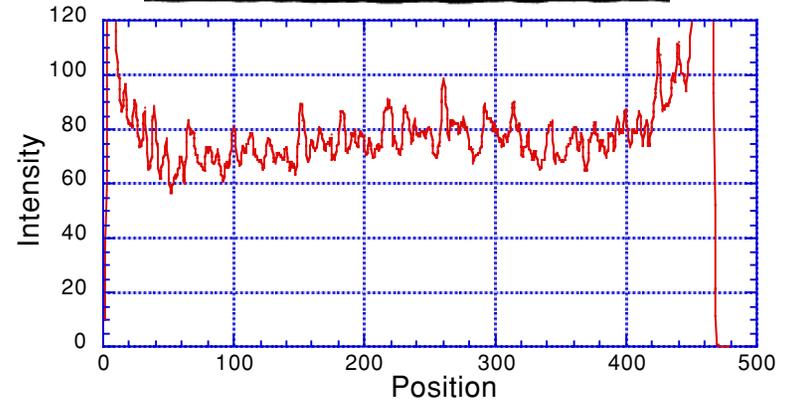
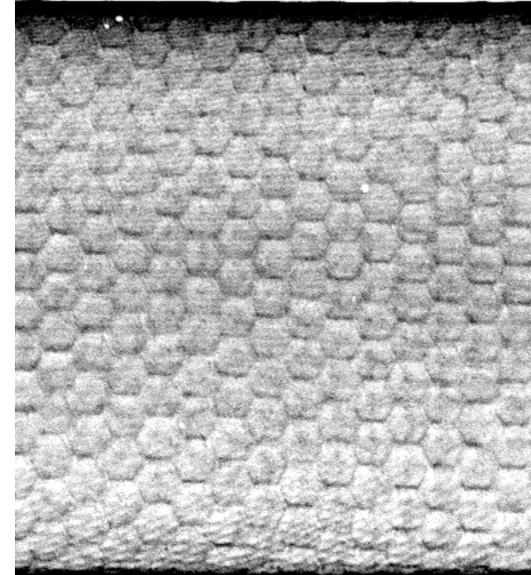
COS BBA#2-B UV full field illumination with Flight TDC



Flight BBA#2-B MCP Stack Fixed Pattern Noise



•Multifiber modulation dominates,



small Moiré at some edges



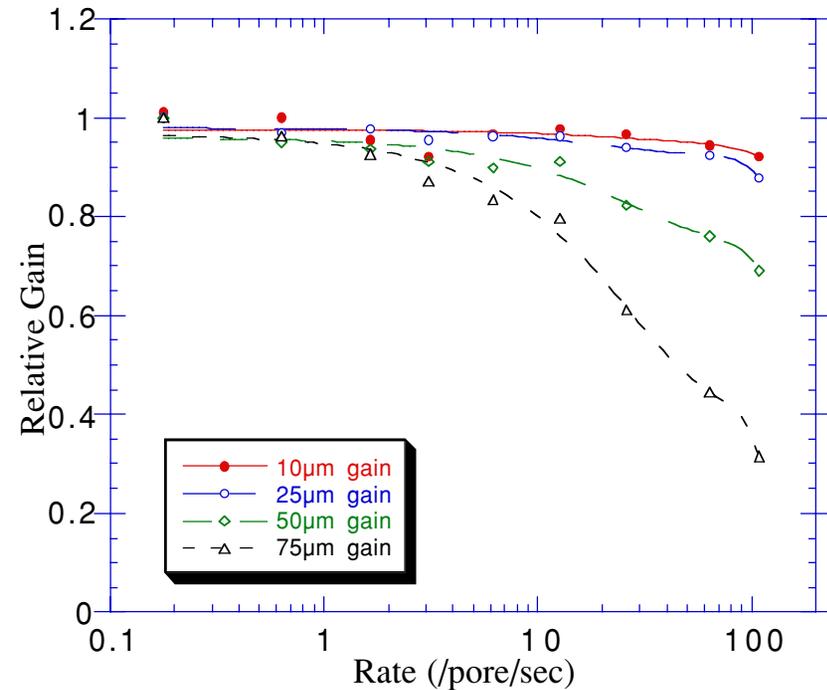
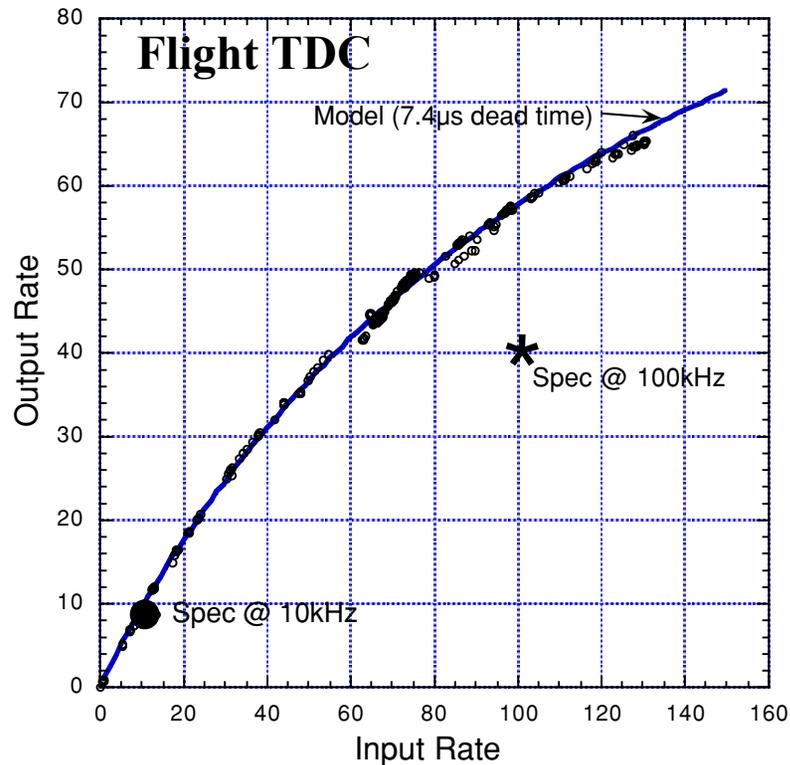
Flat Field Image Stability Test



- A flat field will be taken over at least an area of 80 mm x 0.5mm with 14100 cts per resolution element (35 μ m x 200 μ m)
- The backplate and TDC will be raised in temperature by ≥ 5 deg. C
- Another flat field will be collected.
- The RMS deviation between these flats must not exceed 1% after correcting for thermal distortions using the centroids of the electronic stims.

- Requirements

- Global Output rates of $>40,000$ cts s^{-1} per segment output
- Local Input rates of >5 cts s^{-1} per microchannel
- Deadtime $<10\%$ at input rate of $10,000$ cts s^{-1} & $<60\%$ at $100,000$ cts s^{-1}





QDE Calibration



- DQE specified at 6 wavelengths:

Wavelength (Å)	SoR DQE (%)	Measured (%)
1152	44	49
1216	32	36
1335	25	26
1463	19	21
1560	17	16
1710	11	10

- Measured *after* CsI photocathode deposition
- Absolute calibration measured in center of detector and traceable to NIST calibrated photodiode
- Measured at input angle appropriate for wavelength
- Relative DQE vs. X position performed with detector translation
- Test setup is a flight BBA with QE enhancement grid in place on a test detector.

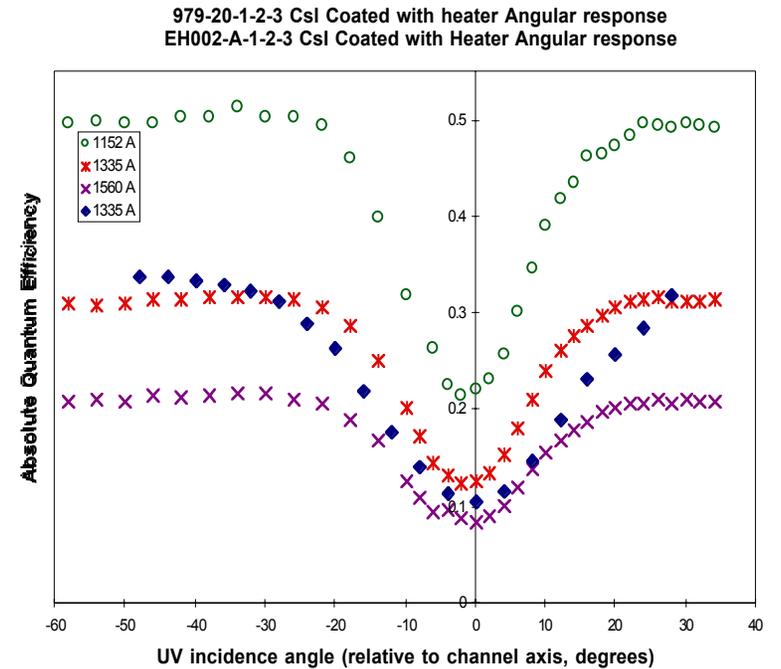
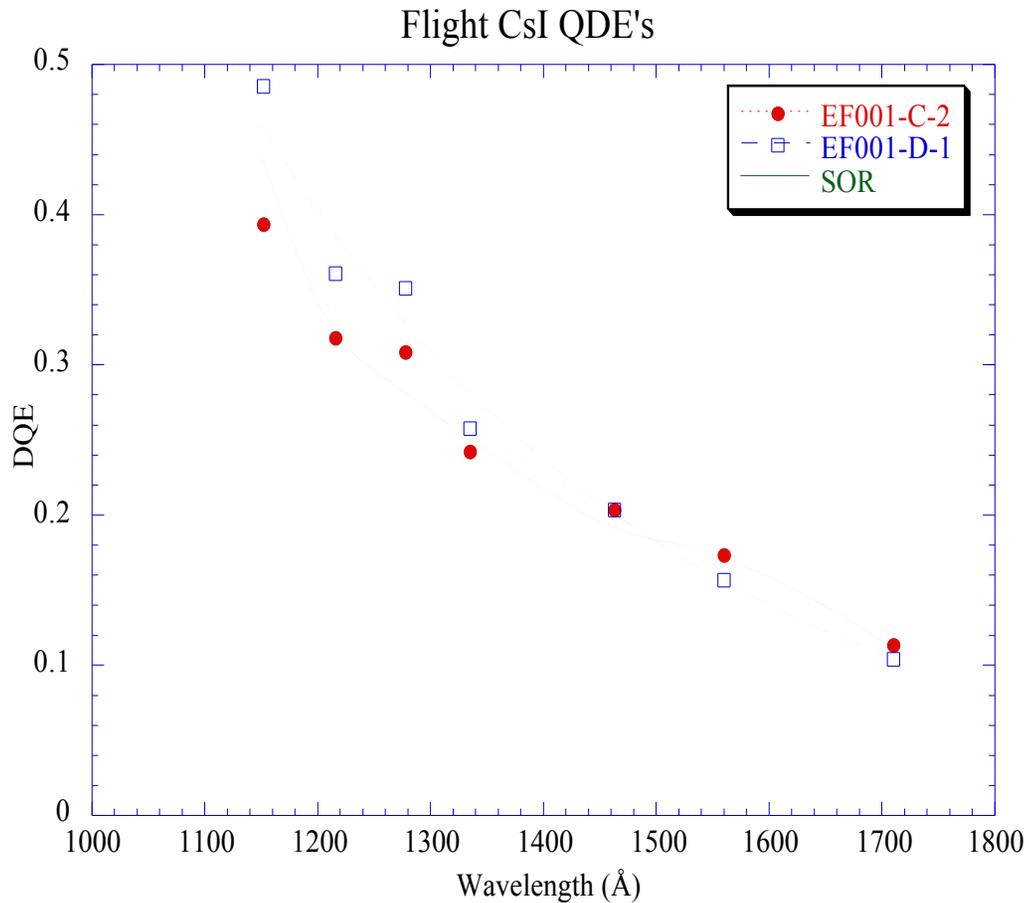


QDE Calibration (cont'd)



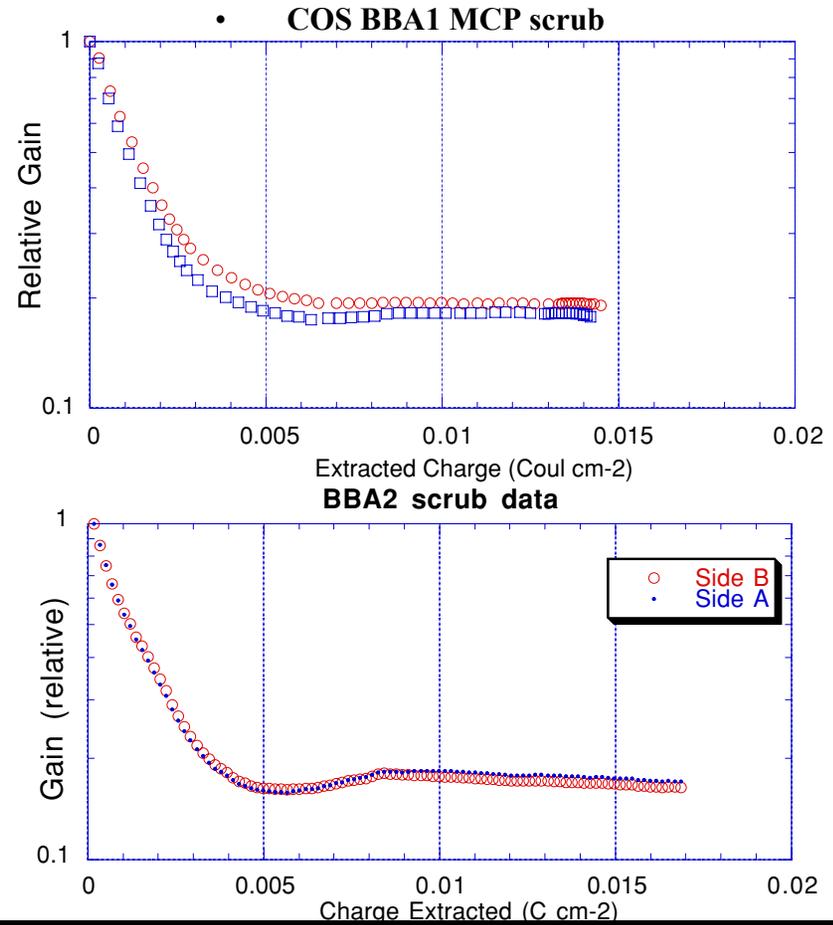
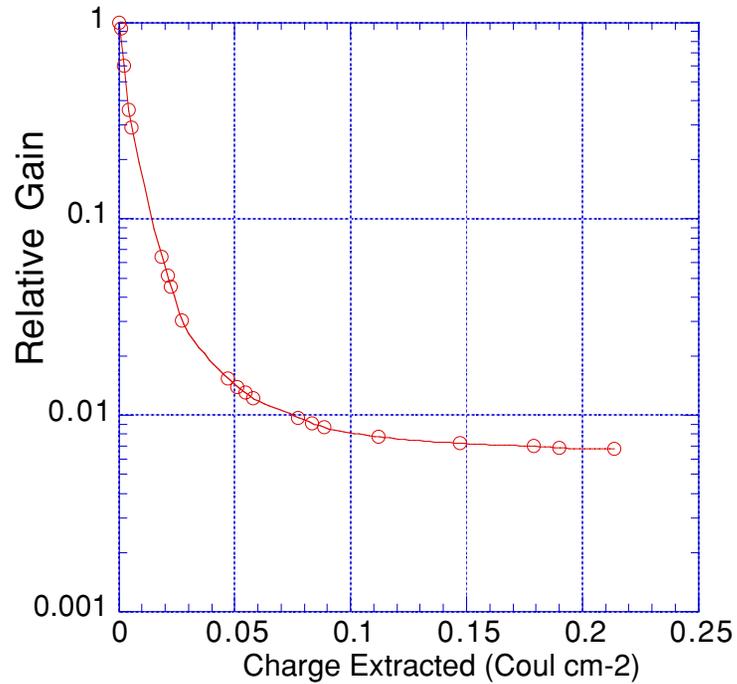
- Grid Shadows
 - measured with collimated light flat field.
 - Effect of f/24 input beam then determined by analysis
- Visible Light Rejection
 - Use two monochromatic He-Ne laser lines (632nm and 543 nm)
 - Calibrated optical (Si) photodiode to measure beam flux
 - All previous tests show that visible QDE to be orders of magnitude below requirement and therefore do not need to characterize QDE curve in visible regime

- Sample QDE plot



Both pre-scrubs on BBA's #1 and #2 show gain stabilization

- COS sample MCP scrub





BBA to Backplate Metrology Plan / Status



- To maintain focus, the MCP's input surface must match the optical focal surface ($R=826$ mm) to ± 100 μm (SoR 7.1.1)
- Since there is ± 0.5 mm of focus adjustment in X, Y, and Z the primary constraint is on the knowledge of the orientation of the MCP surface relative to the backplate surface.
- A CMM with a camera and microscope objective will be used to measure the MCP surface relative to the backplate surface and optical cube to ± 30 μm .
- The BBA will be shimmed such that the vertex of the MCP radius will be 10.4 mm above the backplate surface.
- The flight and spare detectors will be shimmed to within ± 50 μm of the same height.
- The MCP surfaces will be characterized in about 40 locations per segment in two stripes in Y once final MCP and BBA shimming has been performed.
- Procedures are in place (UCB-COS-PRO-1132 and UCB-COS-PRO-1133) and measurements are in progress.